



Register Number:
DATE:

ST. JOSEPH'S COLLEGE (AUTONOMOUS), BANGALORE-27

M.Sc. PHYSICS - III SEMESTER

SEMESTER EXAMINATION- OCTOBER 2019

For supplementary candidates only.

PH 9115 - QUANTUM MECHANICS II

Time-2 1/2 hrs.

Maximum Marks-70

This question paper has 4 printed pages and 2 parts and a table of constants and integrals

PART A

Answer any FIVE full questions.

(5x10=50)

1. Consider a system made up of two identical particles a and b each located at positions \vec{r}_1 and \vec{r}_2 respectively and having wavefunctions: $\psi_a(\vec{r}_1)$ and $\psi_b(\vec{r}_2)$ respectively. Develop the composite wavefunction for the system if the particles are indistinguishable bosons.

2.

- (a) Using the Variational Principle, obtain the estimate of the ground state energy of a particle in a one dimensional box. Assume the wavefunction to have a form of: $\psi_t = A e^{-bx^2}$

(5 Marks)

- (b) The WKB approximation with the connection formulae result in the following expression for

the wave function:
$$\psi(x) = \begin{cases} \frac{D}{\sqrt{p(x)}} \sin \frac{1}{\hbar} \left(\int_x^{x_2} p(x') dx' + \frac{\pi}{4} \right) & x < x_2 \\ \frac{D}{\sqrt{|p(x)|}} e^{-\frac{1}{\hbar} \left(\int_{x_2}^x |p(x')| dx' \right)} & x > x_2 \end{cases}$$
 . Obtain the

energy of the bound states (energy levels) of a particle in a box with a hard wall at $x=0$ and a sloping floor potential of the form $V(x) = \begin{cases} kx & \text{if } x > 0 \\ \infty & \text{otherwise} \end{cases}$ where

k is a positive constant.

(5 Marks)

3. Explain perturbation theory and obtain the first order corrections to the energy for a system having its state space described by non-degenerate eigenstates.
4. Obtain the determinant (for first order correction to the energy) of a system having a three fold degenerate energy level.
5.
 - (a) Starting with the time dependent Schrodinger Equation:

$$i\hbar \frac{\partial}{\partial t} \Psi(\vec{r}, t) = H(\vec{r}, t) \Psi(\vec{r}, t) ,$$
 using separation of variables $\Psi(\vec{r}, t) = \phi(\vec{r}) \psi(t)$, obtain the Schrodinger equation for time dependent perturbation of the type: $H(t) = H_0 + \lambda \hat{W}(t)$.
 - (b) If we assume that at time $t=0$, $|\psi(t=0)\rangle = |\phi_i\rangle$, and that for any time $t \neq 0$ we assume $|\psi(t)\rangle = \sum_n c_n(t) |\phi_n\rangle$, what will the Schrodinger equation reduce to?
6.
 - (a) Explain scattering by a potential. What are the assumptions used in order to develop the scattering theory? **(4 Marks)**
 - (b) What is scattering cross section? Write down the expression for differential cross section and work out the asymptotic form for the scattered wave. **(6 Marks)**
7. Derive the Hamiltonian of a system of two particles in a potential (for the scattering problem) dependent on the distance between the two in the center of mass frame. Express the Hamiltonian in reduced mass, and explain all terms.

PART B

Answer any **FOUR** full questions.

(4x5=20)

8. Two noninteracting particles of mass m are in a one dimensional infinite potential well. Remember, the wavefunction and energies for one particle in an infinite potential are given as:

$$\psi_n(x) = \sqrt{\frac{2}{L}} \sin \frac{n\pi x}{L} \quad \text{and} \quad E_n = \frac{\hbar^2 n^2 \pi^2}{2m L^2} .$$
 What are the wavefunction and energies of the two particles if:
 - a) both are in the ground state? What are these for Fermions and Bosons respectively? **(2 Marks)**
 - b) one is in the ground state and the other in the first excited state? What are these for Fermions and Bosons? **(3 Marks)**
9. Using the Variational Principle, estimate the ground state energy of a Simple Harmonic Oscillator (with potential $V(x) = \frac{1}{2} m \omega^2 x^2$). Use a trial wavefunction of the form

$$\psi_t(x) = \frac{A}{x^2 + b} .$$
 You may take the expectation value of the Kinetic Energy to be given as:

$$\langle T \rangle = \frac{\hbar^2}{2m} \frac{1}{2b} .$$
10. A particle is trapped in a well. External to this well, the potential has the form $V(x) = -kx$. Draw a figure, clearly demarcating the turning points. Obtain the expression for the probability of the particle tunneling out of the potential.
11. Calculate the first order correction to $E_3^{(0)}$ for a particle in a one-dimensional box with walls

at $x=0$ and $x=L$ due to a perturbation: $W = \lambda \hat{W} = 10^{-3} E_1 \frac{x}{L}$.

12. We know that for the time dependent perturbation, the component of the state $|\phi_k\rangle$ to which the state $|\phi_i\rangle$ transitions to is given to a first order as:

$$b_k^{(1)}(t) = \frac{1}{i\hbar} \int_0^t e^{i\omega_{ki}t'} \langle \phi_k | \hat{W} | \phi_i \rangle dt' .$$

Furthermore, the transition probability from an initial state $|\phi_i\rangle$ to a final state $|\phi_f\rangle$ can be deduced from this as $P_{if} = \lambda^2 |b_f^{(1)}(t)|^2$. For $\hat{W}(t) = \hat{W}_0$ (a constant) compute P_{if} .

13. Use the Born approximation ($f_k^{(B)}(\theta, \phi) = \frac{2\mu}{\hbar^2 K} \int r' V(r') \sin(Kr') dr'$) to calculate the differential cross-section (remember, that the expression above is the scattering amplitude) for a central potential of the form $V(\vec{r}) = \beta \frac{e^{-\alpha r}}{r}$ (which is the Yukawa potential).

(Some) Physical Constants

[Constants: $h = 6.626070 \times 10^{-34}$ J s (**Planck's constant**), $1\text{eV} = 1.6 \times 10^{-19}$ J (**electron volt to Joules**), $c = 2.99792458 \times 10^8$ m/s (**speed of light**), $1\text{\AA} = 1 \times 10^{-10}$ m (**Angstrom to meters**), $e = 1.602176 \times 10^{-19}$ C (**electronic charge**), $\epsilon_0 = 8.85418782 \times 10^{-12}$ m³kg⁻¹s⁴A² (**permittivity of free space**), $m_{\text{proton}} = 1.672621898 \times 10^{-27}$ kg (**mass of proton**), $m_{\text{electron}} = 9.10938356 \times 10^{-31}$ kg (**mass of electron**), $m_{\text{neutron}} = 1.674927471 \times 10^{-27}$ kg (**mass of neutron**), $a = 5.029 \times 10^{-10}$ m (**Bohr radius**), $\alpha = 1/137$ (**Fine Structure Constant**), $G = 6.674 \times 10^{-11}$ m³kg⁻¹s⁻² (**Gravitational constant**), $M_{\odot} = 1.9891 \times 10^{30}$ kg (**Solar mass**), $R_{\odot} = 6.9 \times 10^8$ m (**Sun's Radius**), $\sigma = 5.67 \times 10^{-8}$ W m⁻² K⁻⁴ (**Stefan-Boltzmann constant**), $M_{\text{Earth}} = 5.97 \times 10^{27}$ kg (**Mass of Earth**), $D_{\text{earth-sun}} = 1.49 \times 10^{11}$ m (**Earth-Sun distance**), 1 inch = 2.54 cm, 1 foot = 12 inches]

(a) Table of (some) Integrals

Gamma Function:
$\Gamma(x) = \int_0^{\infty} t^{x-1} e^{-t} dt$
$\Gamma(n) = (n-1)!$
$\Gamma\left(\frac{1}{2} + n\right) = \frac{(2n)!}{4^n n!} \sqrt{\pi}$

(a) $\int_0^{\infty} e^{-bt} dt = \frac{1}{b}$

(k) $\int \frac{t^2}{(t^2+b^2)^2} dt = \left(-\frac{t}{(2b^2+2t^2)} + \frac{1}{2b} \tan^{-1}\left(\frac{t}{b}\right) \right)$

(b) $\int_0^{\infty} t e^{-2bt} dt = \frac{1}{4b^2}$

(l) $\int \frac{1}{(t^2+b^2)^3} dt = \frac{3}{8b^5} \left(\frac{5/3 b^3 t + b t^3}{(b^2+t^2)^2} + \tan^{-1}\left(\frac{t}{b}\right) \right)$

$$(c) \int_0^{\infty} t^2 e^{-2bt} dt = \frac{1}{4b^3}$$

$$(d) \int_0^{\infty} t^3 e^{-2bt} dt = \frac{3}{8b^4}$$

$$(e) \int_0^{\infty} t^4 e^{-2bt} dt = \frac{3}{4b^5}$$

$$(f) \int_0^{\infty} t^5 e^{-2bt} dt = \frac{15}{8b^6}$$

$$(g) \int_0^{\infty} t^6 e^{-2bt} dt = \frac{45}{8b^7}$$

$$(h) \int \frac{1}{t^2+b^2} dt = \frac{1}{b} \tan^{-1}\left(\frac{t}{b}\right)$$

$$(i) \int \frac{1}{(t^2+b^2)^2} dt = \frac{1}{2b^3} \left(\frac{bt}{(b^2+t^2)} + \tan^{-1}\left(\frac{t}{b}\right) \right)$$

$$(j) \int \frac{1}{(t^2+b^2)^4} dt = \frac{1}{16b^7} \left(\frac{15t^5b+40b^3t^3+33b^5t}{(3t^6+9bt^4+9b^3t^2+b^5)} + 5 \tan^{-1}\left(\frac{t}{b}\right) \right)$$

$$(m) \int \frac{t^2}{(t^2+b^2)^4} dt = \frac{1}{16b^5} \left(\frac{bt^5+8/3b^3t^3-3b^5t}{(b^2+t^2)^3} + \tan^{-1}\left(\frac{t}{b}\right) \right)$$

$$(n) \int \sqrt{a/x-1} dx = x\sqrt{a/x-1} + a \tan^{-1}(\sqrt{a/x-1})$$

$$(o) \int \sqrt{1-ax} dx = -\frac{2(1-ax)^{3/2}}{3a}$$

$$(p) \int \sqrt{1-ax^2} dx = \frac{1}{2} x \sqrt{1-ax^2} + \frac{\sin^{-1}\sqrt{ax}}{2\sqrt{a}}$$

$$(q) \int_{-\infty}^{\infty} e^{-\alpha t^2+i\omega t} dt = \sqrt{\frac{\pi}{\alpha}} e^{-\frac{\omega^2}{4\alpha}}$$

$$(r) \int_0^{\infty} t^4 e^{-\alpha^2 t^2} dt = \frac{3\sqrt{\pi}}{8\alpha^5}$$

$$(s) \int_0^{\infty} t^n e^{-st} dt = \frac{n!}{s^{n+1}} \quad (\text{Laplace Transform})$$